## APPLICATIONS OF MODAL METHODS

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DAMTP
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## ESTIMATORS

For the bispectrum the estimator takes the general form

We can put this in a general form by defining

$$
\begin{gathered}
\left\langle\mathfrak{a}_{\wp>}\right\rangle \equiv\left\langle a_{l_{1} m_{1}} a_{l_{2} m_{2}} \ldots a_{l_{p} m_{p}}\right\rangle \\
\mathfrak{C}_{\wp \wp \wp^{\prime}}^{-1} \equiv C_{l_{1} m_{1}, l_{1}^{\prime} m_{1}^{\prime}}^{-\ldots C_{l_{p} m_{p}, l_{p}^{\prime} m_{2}^{\prime} p}^{-1}}
\end{gathered}
$$

Where $\wp$ represents the $\wp=\left\{l_{1}, m_{1}, l_{2}, m_{2}, \ldots, l_{p}, m_{p}\right\}$ degrees of freedom

## ESTIMATORS

The estimator for a general polyspectrum is then defined as

$$
\overline{\mathcal{E}} \equiv \frac{\sum_{\wp_{\wp} \wp^{\prime}}\left\langle\mathfrak{a}_{\wp}\right\rangle \mathfrak{C}_{\wp 申 \wp^{\prime}}^{-1}\left(\mathfrak{a}_{\wp}-\mathfrak{a}_{\wp}^{\text {lin }}\right)}{\sum_{\wp \wp \wp^{\prime}}\left\langle\mathfrak{a}_{\wp}\right\rangle \mathfrak{C}_{\wp \wp^{\prime}}^{-1}\left\langle\mathfrak{a}_{\wp}\right\rangle}
$$

where $\mathfrak{a}_{\wp}^{\text {lin }}$ is the appropriate linear term

## ESTIMATORS

We will now go one step further by defining the weighted vectors (and matrix)

$$
\mathcal{A}_{\wp}=\frac{\left\langle\mathfrak{a}_{\wp}\right\rangle}{\sqrt{C_{l_{1}} C_{l_{2}} \ldots C_{l_{p}}}}, \quad \mathcal{B}_{\wp}=\frac{\mathfrak{a}_{\wp}-\mathfrak{a}_{\wp}^{l i n}}{\sqrt{C_{l_{1}} C_{l_{2}} \ldots C_{l_{p}}}}, \quad \mathcal{C}_{\wp \wp^{\prime}}=\frac{\mathfrak{C}_{\wp \wp^{\prime}}}{\sqrt{C_{l_{1}} C_{l_{1}^{\prime}} \ldots C_{l_{p}} C_{l_{p}^{\prime}}}},
$$

And we can then write the estimator in matrix form as

$$
\overline{\mathcal{E}}=\frac{\mathcal{A}^{T} \mathcal{C}^{-1} \mathcal{B}}{\mathcal{A}^{T} \mathcal{C}^{-1} \mathcal{A}}
$$

## BASIS

If we then suppose the existence of an orthonormal basis

$$
\sum_{\wp} \mathcal{R}_{n_{\wp}} \mathcal{R}_{n^{\prime} \wp}=\delta_{n n^{\prime}} \quad\left(\mathcal{R} \mathcal{R}^{T}=I\right)
$$

built from some separable functions $\mathcal{R}=\lambda \mathcal{Q}$

$$
\begin{gathered}
\mathcal{R}=\frac{\mathcal{G}_{m_{1} m_{2} m_{3}}^{l_{1} l_{2} l_{3}}}{v_{l_{1}} v_{l_{2}} v_{l_{3}}} R_{n l_{1} l_{2} l_{3}} \\
R_{n l_{1} l_{2} l_{3}}=\lambda_{n m} Q_{n l_{1} l_{2} l_{3}}\left(=q_{i} q_{j} q_{k}+5 \mathrm{perms}\right)
\end{gathered}
$$

## BASIS

Then we can decompose our theory representing it as a set of modal coefficients

$$
\begin{gathered}
\mathcal{A}_{\wp}=\sum_{n} \alpha_{n} R_{n \wp} \quad\left(\mathcal{A}=\mathcal{R}^{T} \alpha\right) \\
\alpha=\mathcal{R A}
\end{gathered}
$$

We will truncate our basis at some nmax so so we can also define a projection operator

$$
\mathcal{P}=\mathcal{R}^{T} \mathcal{R}
$$

And we take our theory to be completely described by this basis

$$
\mathcal{P} \mathcal{A}=\mathcal{A}
$$

## SIMULATION

This method can be used to simulate maps with a given bispectrum and trispectrum

$$
\begin{gathered}
a_{l m}=a_{l m}^{G}+\frac{1}{6} F_{N L} a_{l m}^{B}+\frac{1}{24} G_{N L} a_{l m}^{T} \\
a_{l m}^{B}=\sum_{l_{i} m_{i}} \int Y_{l_{1} m_{1}} Y_{l_{2} m_{2}} Y_{l_{3} m_{3}} b_{l_{1} l_{2} l_{3}} \frac{a_{l_{2} m_{2}}^{G}}{C_{l_{2}}} \frac{a_{l_{3} m_{3}}^{G}}{C_{l_{3}}} \\
a_{l m}^{T}=\sum_{l_{i} m_{i}} \int Y_{l_{1} m_{1}} Y_{l_{2} m_{2}} Y_{l_{3} m_{3}} Y_{l_{4} m_{4}} t_{l_{1} l_{2} l_{3} l_{4}} \frac{a_{l_{2} m_{2}}^{G}}{C_{l_{2}}} \frac{a_{l_{3} m_{3}}^{G}}{C_{l_{3}}} \frac{a_{l_{4} m_{4}}^{G}}{C_{l_{4}}}
\end{gathered}
$$

## SIMULATION

Using the expansion the nonGaussian contributions can be easily calculated

$$
\begin{aligned}
a_{l m}^{B} & =\sum_{n} \bar{\alpha}_{n}^{Q} \frac{q_{l}^{\{i}}{v_{l} \sqrt{C_{l}}} \int d^{2} \hat{\mathbf{n}} Y_{l m}(\hat{\mathbf{n}}) M^{j}(\hat{\mathbf{n}}) M^{k\}}(\hat{\mathbf{n}}) \\
M^{i}(\hat{\mathbf{n}}) & =\sum_{l m} \frac{q_{l}^{i} Y_{l m}(\hat{\mathbf{n}}) a_{l m}^{G}}{v_{l} \sqrt{C_{l}}}
\end{aligned}
$$



## SIMULATION

To test the accuracy of the method we simulated maps using both the primordial and CMB decompositions and then applied both


Map the primordial and CMB estimators to both sets to produce consistent results

|  | Ideal simulations |  | WMAP5 simulations |  |
| :---: | :---: | :---: | :---: | :---: |
|  | Average | St. Dev. | Average | St. Dev. |
| Primordial estimator | 292.9 | 104.8 | 297.7 | 152.1 |
| Late-time estimator | 300.6 | 104.9 | 278.7 | 160 |
| Internal st. dev. | 38.5 |  | 102.6 |  |



## BASIS

We can perform the same modal decomposition on the data and the covariance

$$
\begin{aligned}
\alpha & =\mathcal{R} \mathcal{A} \\
\beta & =\mathcal{R B} \longrightarrow \mathcal{P B}=\mathcal{R}^{T} \beta \\
\zeta & =\mathcal{R C} \mathcal{R}^{T} \\
\mathcal{E} & \equiv \frac{\alpha^{T} \zeta^{-1} \beta}{\alpha^{T} \zeta^{-1} \alpha} \\
& =\frac{(\mathcal{R A})^{T} \mathcal{R C} \mathcal{C}^{-1} \mathcal{R}^{T} \mathcal{R B}}{\mathcal{R} \mathcal{A}^{T} \mathcal{R} \mathcal{C}^{-1} \mathcal{R}^{T} \mathcal{R A}}=\frac{\mathcal{A}^{T} \mathcal{P} \mathcal{C}^{-1} \mathcal{P B}}{\mathcal{A}^{T} \mathcal{P} \mathcal{C}^{-1} \mathcal{P A}}
\end{aligned}
$$

## BASIS

We can understand the effect of the projection by considering

$$
\begin{aligned}
\mathcal{A}=\left[\begin{array}{c}
\mathcal{A}_{\|} \\
0
\end{array}\right] \quad \mathcal{B} & =\left[\begin{array}{c}
\mathcal{B}_{\|} \\
\mathcal{B}_{\perp}
\end{array}\right] \quad \mathcal{C}^{-1}=\left[\begin{array}{cc}
\mathcal{C}_{\|}^{-1} & \mathcal{C}_{\times}^{-1} \\
\mathcal{C}^{-1}{ }_{\times} & \mathcal{C}_{\perp}^{-1}
\end{array}\right] \\
\mathcal{X}_{\|} & \equiv \mathcal{P} \mathcal{X} \\
\mathcal{X}_{\perp} & \equiv(I-\mathcal{P}) \mathcal{X} \\
\mathcal{M}_{\|} & \equiv \mathcal{P} \mathcal{M} \mathcal{P} \\
\mathcal{M}_{\perp} & \equiv(I-\mathcal{P}) \mathcal{M}(I-\mathcal{P}) \\
\mathcal{M}_{\times} & \equiv \mathcal{P} \mathcal{M}(I-\mathcal{P})
\end{aligned}
$$

## BASIS

We can understand the effect of the projection by

$$
\begin{gathered}
\overline{\mathcal{E}}=\frac{\mathcal{A}_{\|}\left(\mathcal{C}_{\|}^{-1} \mathcal{B}_{\|}+\mathcal{C}_{\times}^{-1} \mathcal{B}_{\perp}\right)}{\mathcal{A}_{\|}^{T} \mathcal{C}_{\|}^{-1} \mathcal{A}_{\|}} \\
\mathcal{E}=\frac{\mathcal{A}_{\|} \mathcal{C}_{\|}^{-1} \mathcal{B}_{\|}}{\mathcal{A}_{\|}^{T} \mathcal{C}_{\|}^{-1} \mathcal{A}_{\|}}
\end{gathered}
$$

The difference is the projection of contamination from the orthogonal space into the subspace

## INVERSE COVARIANCE

## Can we even calculate the covariance in the modal space? Yes!

$$
\begin{aligned}
& \left\langle\beta_{n} \beta_{n^{\prime}}\right\rangle=\sum_{l_{i} m_{i} l_{i}^{\prime} m_{i}^{\prime}}\left\langle\left(\frac{\mathcal{G}_{m_{1} m_{2} m_{3}}^{l_{1} l_{2} l_{3}}}{v_{l_{1}} v_{l_{2}} v_{l_{3}}} \frac{a_{l_{1} m_{1}} a_{l_{2} m_{2}} a_{l_{3} m_{3}}-3 C_{l_{1} m_{1}, l_{2} m_{2}} a_{l_{3} m_{3}}}{\sqrt{C_{l_{1}} C_{l_{2}} C_{l_{3}}}} R_{n l_{1} l_{2} l_{3}}\right)\right. \\
& \left.\times\left(\frac{\mathcal{G}_{m_{1}^{\prime} m_{2}^{\prime} m_{3}^{\prime}}^{l_{1}^{\prime} l_{l}^{\prime} l_{3}^{\prime}}}{v_{l_{1}^{\prime}} v_{l_{2}^{\prime}} v_{l_{3}^{\prime}}} \frac{a_{l_{1}^{\prime} m_{1}^{\prime}} a_{l_{2}^{\prime} m_{2}^{\prime}} a_{l_{3}^{\prime} m_{3}^{\prime}}-3 C_{l_{1}^{\prime} m_{1}^{\prime}, l_{2}^{\prime} m_{2}^{\prime}} a_{l_{3}^{\prime} m_{3}^{\prime}}}{\sqrt{C_{l_{1}^{\prime}} C_{l_{2}^{\prime}} C_{l_{3}^{\prime}}}} R_{n l_{1}^{\prime} l_{2}^{\prime} l_{3}^{\prime}}\right)\right\rangle \\
& =\sum_{l_{i} m_{i} l_{i}^{\prime} m_{i}^{\prime}} \frac{\mathcal{G}_{m_{1} m_{2} m_{3}}^{l_{1} l_{2} l_{3}} \mathcal{G}_{m_{1}^{\prime} m_{2}^{\prime} m_{3}^{\prime}}^{l_{1}^{\prime} l_{2}^{\prime} l_{3}^{\prime}}}{v_{l_{1}} v_{l_{2}} v_{l_{3}} v_{l_{1}^{\prime}} v_{l_{2}^{\prime}} v_{l_{3}^{\prime}}} R_{n l_{1} l_{2} l_{3}}\left[6\left\langle a_{l_{1} m_{1}} a_{l_{1}^{\prime} m_{1}^{\prime}}\right\rangle\left\langle a_{l_{2} m_{2}} a_{l_{2}^{\prime} m_{2}^{\prime}}\right\rangle\left\langle a_{l_{3} m_{3}} a_{l_{3}^{\prime} m_{3}^{\prime}}\right\rangle\right. \\
& +9\left\langle a_{l_{1} m_{1}} a_{l_{2} m_{2}}\right\rangle\left\langle a_{l_{1}^{\prime} m_{1}^{\prime}} a_{l_{2}^{\prime} m_{2}^{\prime}}\right\rangle\left\langle a_{l_{3} m_{3}} a_{l_{3}^{\prime} m_{3}^{\prime}}\right\rangle-9 C_{l_{1} m_{1}, l_{2} m_{2}}\left\langle a_{l_{1}^{\prime} m_{1}^{\prime}} a_{l_{2}^{\prime} m_{2}^{\prime}}\right\rangle\left\langle a_{l_{3} m_{3}} a_{l_{3}^{\prime} m_{3}^{\prime}}\right\rangle \\
& \left.-9\left\langle a_{l_{1} m_{1}} a_{l_{2} m_{2}}\right\rangle C_{l_{1}^{\prime} m_{1}^{\prime}, l_{2}^{\prime} m_{2}^{\prime}}\left\langle a_{l_{3} m_{3}} a_{l_{3}^{\prime} m_{3}^{\prime}}\right\rangle+9 C_{l_{1} m_{1}, l_{2} m_{2}} C_{l_{1}^{\prime} m_{1}^{\prime}, l_{2}^{\prime} m_{2}^{\prime}}\left\langle a_{l_{3} m_{3}} a_{l_{3}^{\prime} m_{3}^{\prime}}\right\rangle+\ldots\right] R_{n l_{1}^{\prime} l_{2}^{\prime} l_{3}^{\prime}} \\
& =6 \sum_{l_{i} m_{i} l_{i}^{\prime} m_{i}^{\prime}} \frac{\mathcal{G}_{m_{1} m_{2} m_{3}}^{l_{1} l_{2} l_{3}} \mathcal{G}_{m_{1}^{\prime} m_{2}^{\prime} m_{3}^{\prime}}^{l_{1}^{\prime} l_{2}^{\prime} l_{3}^{\prime}}}{v_{l_{1}} v_{l_{2}} v_{l_{3}} v_{l_{1}^{\prime}} v_{l_{2}^{\prime}} v_{l_{3}^{\prime}}} R_{n l_{1} l_{2} l_{3}} \frac{C_{l_{1} m_{1}, l_{1}^{\prime} m_{1}^{\prime}} C_{l_{2} m_{2}, l_{2}^{\prime} m_{2}^{\prime}} C_{l_{3} m_{3}, l_{3}^{\prime} m_{3}^{\prime}}}{\sqrt{C_{l_{1}} C_{l_{2}} C_{l_{3}} C_{l_{1}^{\prime}} C_{l_{2}^{\prime}} C_{l_{3}^{\prime}}}} R_{n^{\prime} l_{1}^{\prime} l_{2}^{\prime} l_{3}^{\prime}}
\end{aligned}
$$

## INVERSE COVARIANCE

Also as all covariance matrices are symmetric positive definite they have a Cholesky decomposition

$$
\zeta=\tilde{\lambda} \tilde{\lambda}^{T}
$$

And we can absorb the covariance into our modes. This amounts to a re-orthogonalisation to an uncorrelated orthonormal basis

$$
\begin{gathered}
\alpha^{\prime}=\tilde{\lambda}^{-1} \alpha \quad \beta^{\prime}=\tilde{\lambda}^{-1} \beta \\
\mathcal{E}=\frac{\alpha^{\prime T} \beta^{\prime}}{\alpha^{\prime T} \alpha^{\prime}}, \quad \zeta^{\prime}=I
\end{gathered}
$$

## RECONSTRUCTION

We have $\langle\beta\rangle=\alpha$ so can reconstruct the best fit bispectrum to the data by using the $\beta$ as our $\alpha$. If we have constructed a primordial basis as well then we can use the decomposition of projected primordial modes to find the best fit primordial bispectrum



## RECONSTRUCTION

In addition to constraining particular models we can perform a blind search

$$
F_{N L}^{2}=\frac{\beta^{\prime T} \beta^{\prime}}{\alpha^{\prime T} \alpha^{\prime}}
$$



## CONTAMINANTS

As we expect the covariance matrix to be the identity we can use principle component analysis to identify the shape of contaminants.

We first calculate the covariance matrix for beta from simulations

$$
V \zeta V^{T}=D
$$

And then find the rotation which diagonalises
it. This is equivalent to performing an eigen decomposition. The result is that you obtain a new orthonormal basis but now your modes are uncorrelated and ordered from greatest to least variance.


## CONTAMINANTS

WMAP inhomogeneous noise


## CONTAMINANTS

WMAP Mask



## CONTAMINANTS

## Point sources




## CONCLUSIONS



